

1. INTRODUCTION

At GMT 2024-04-26, 117/02:45, the International Space Station (ISS) was to begin about a 6.7-minute reboost using the Progress 87P thrusters. Figure 1 shows the visiting vehicles' layout updated as of 2024-03-15 with the Progress vehicle as it was docked with its thrusters facing aftwards, putting thrust and the necessary orbital mechanics into play so as to speed up the ISS in its direction of flight. This directional acceleration (increase in velocity), resulted in the altitude elevation of just under 1 km for the space station during this dynamic event. An intended ΔV metric of 0.60 m/s for the massive space station was predicted.

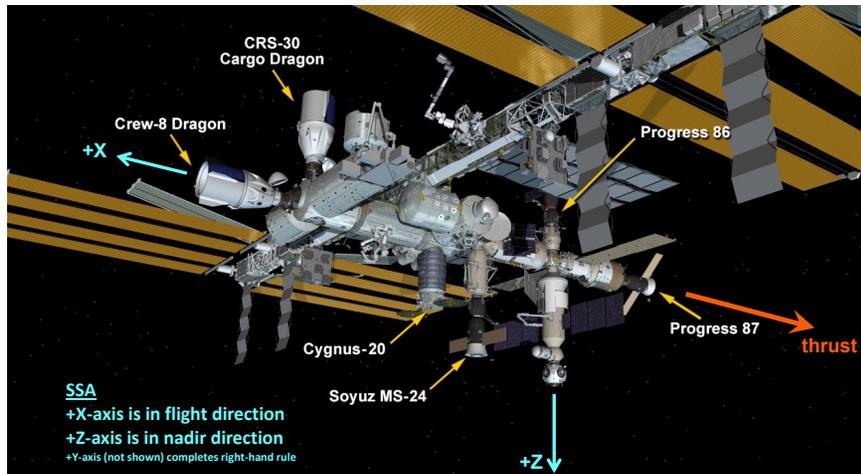


Fig. 1: Progress 87P's Location and Alignment during Reboost.

2. QUALIFY

The information shown in Figure 2 on page 3 was calculated from the Space Acceleration Measurement System (SAMS) sensor 121f08 measurements made in the Columbus module from a sensor mounted on in the COL1A3 rack. This color spectrogram plot shows increased structural vibration excitation contained mostly below 2 Hz or so, and a ~6.9-minute reboost (thruster firing) event was observed as annotated in black. We attribute much of the structural vibration increase to Russian

Segment (RS) attitude control from about GMT 01:50 to about 03:22 as marked with white annotations. The RS thrusters are usually used for station attitude control during the time around the reboost activity. This is expected, and typical behavior. The increased structural vibrations are evident as more noticeable horizontal streaks (structural/spectral peaks) that change from quieter (green/yellow) to more energetic (orange/red) sporadically during this period of RS control. The flare up of these nebulous horizontal (spectral peak) streaks are the tell-tale signatures of large space station appendages as they flex, twist, or bend at their natural frequencies in reaction to impulsive attitude control thruster forces. The actual reboost activity itself lasted ~6.9 minutes as evidenced by slightly more pronounced, vertical orange-red streaks in Figure 2 starting about GMT 02:35:17.

Comparing the same reboost event across all 3 laboratories of the ISS we expect the most energetic vibratory response from the SAMS sensors in the European lab (Columbus module), followed by sensors in the Japanese lab (JEM), and lastly from SAMS sensors in the US lab. This stems from the location of these sensors with respect to structural dynamics of the gargantuan, and quite massive space station. This analysis focused on the power spectral density of vibratory accelerations below 10 Hz, which provides a measure of the intensity of vibratory motion at the natural frequencies of the larger space station structures (e.g. solar array panels and main truss). At higher frequencies (up to 200 Hz), the SAMS sensors usually diverge greatly in terms of acceleration magnitude and frequency components as higher frequency vibrations tend to be more localized, i.e. "mostly" in/around the rack where the sensor is mounted, and due to equipment operations or crew activity in the vicinity.

For science operations and general situational awareness, it is prudent to be aware that the transient and vibratory environment (primarily below about 10 Hz or so) is impacted not only during the relatively brief reboost event itself, but also during the relatively longer span of Russian Segment (RS) attitude control too. The difference being that during the reboost itself, the dominant factor might be considered to be the highly-directional step in the X-axis acceleration, while in the much longer case of RS attitude control, the dominant impact was the excitation of lower-frequency vibrational modes of large space station structures.

3. QUANTIFY

While the spectrogram in the previous “Qualify” section crudely shows acceleration magnitude on a color scale – actually, power spectral density magnitude – we now seek to better quantify the microgravity environment impact of the reboost event across multiple SAMS sensor heads distributed across all 3 main laboratories of the ISS with more intuitive metrics. Figure 3 on page 4 through Figure 7 on page 7 show interval average acceleration results computed from SAMS measurements. Note for all but the es18 sensor, the tell-tale X-axis step that started at GMT 02:35:17 and had a duration of ~6.9 minutes. Information from flight controllers indicated that this reboost event would provide a space station rigid body ΔV of 0.60 meters/second and the SAMS analysis indicated with red annotations in these interval average plots closely match the predicted value. SAMS does not directly measure altitude, but flight controllers indicated that the ISS would gain ~0.96 km in altitude above the Earth as a result of this reboost activity.

Notable Structural Response/Excitation

Beyond what might be considered typical, we note here the spectral “bloom” seen during the reboost span, and most notably in Figure 5 and Figure 6 – for the two sensor heads mounted in the Columbus module – a follow-up per-axis power spectral density plot using spectral averaging from GMT 02:35:45 to 02:41:30 – reveals the majority of that “noise” riding on the reboost is due to a strong spectral peak (a brief spectral “bloom”), ostensibly a strong, structural resonant response at about 0.28 Hz, strongest on the X-, then Z-, then Y-axis. This spectral “bloom” is also conspicuous in the JEM data and least so in the LAB data.

Note that the interval average processing effectively low-pass filtered the data so as to help emphasize the acceleration step that occurs on the X-axis during the reboost event. It should also be noted that we flipped the polarity of each axis (inverted each) in the SAMS plots owing to a polarity inversion issue inherent in SAMS transducers. A somewhat crude quantification of the reboost as measured by these distributed SAMS sensors is also given in Table 1 – expectedly consistent impact results measured by SAMS throughout the space station structure.

Table 1. X-axis steps (mg) during reboost event for multiple SAMS sensors.

Sensor	X-Axis (mg)	Location
121f02	0.148	COL1A1 (ER3)
121f03	0.149	LAB1O1 (ER2)
121f05	0.148	JPM1F1 (ER5)
121f08	0.148	COL1A3 (EPM)
es18	0.150	MSRR (ER6)

4. CONCLUSION

The SAMS measurements for multiple sensor heads distributed across all 3 main labs of the ISS was analyzed and showed an **X-axis step during the Progress 87P reboost of about 0.15 mg**. Furthermore, calculations based on SAMS sensors indicate a ΔV metric of just about 0.60 m/s was achieved, matching the value that flight controllers had predicted.

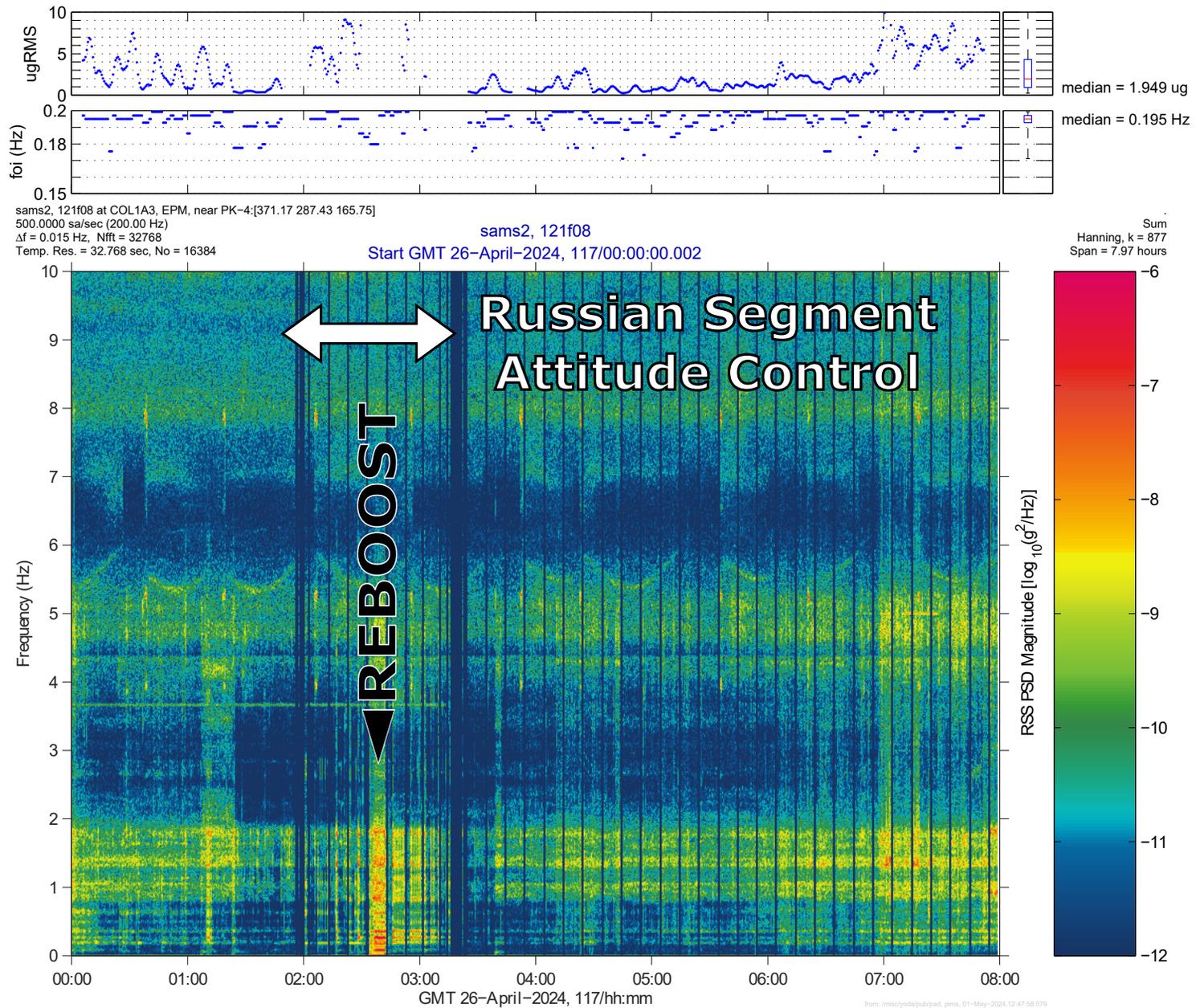


Fig. 2: 10 Hz Spectrogram showing Progress 87P Reboost on GMT 2024-04-26 from a SAMS Sensor in the Columbus module.

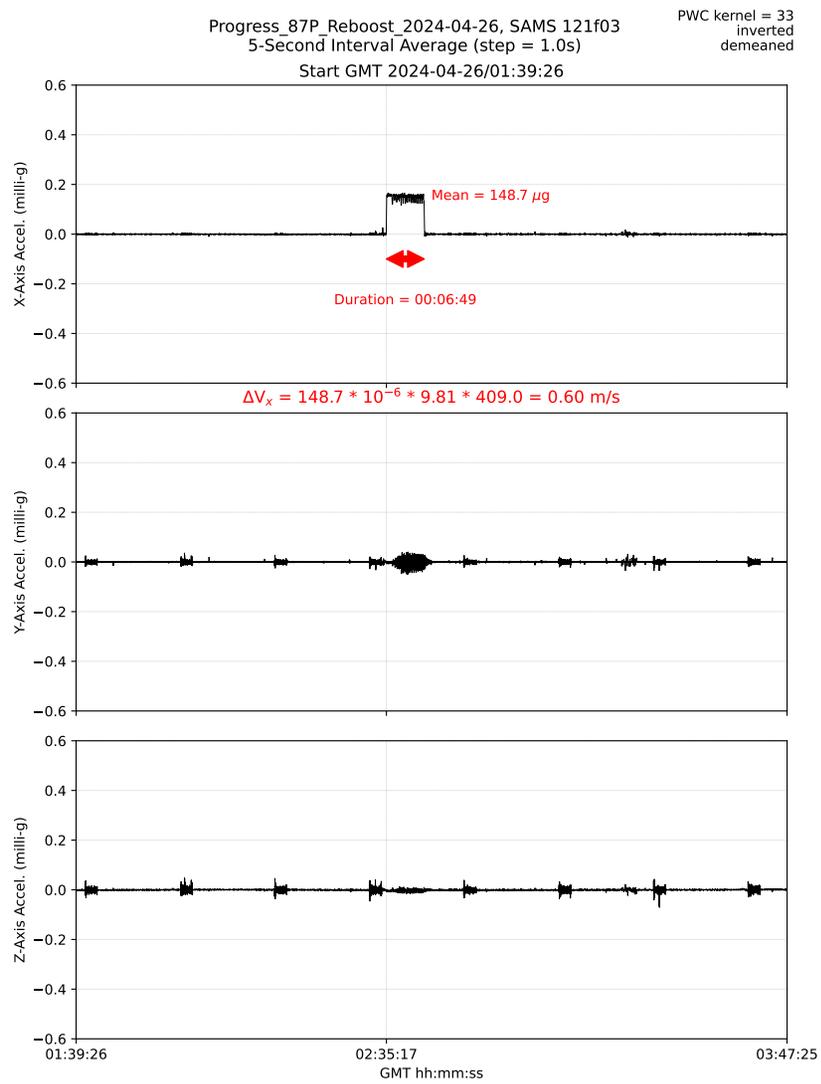


Fig. 3: 5-sec interval average for SAMS 121f03 sensor in the LAB.

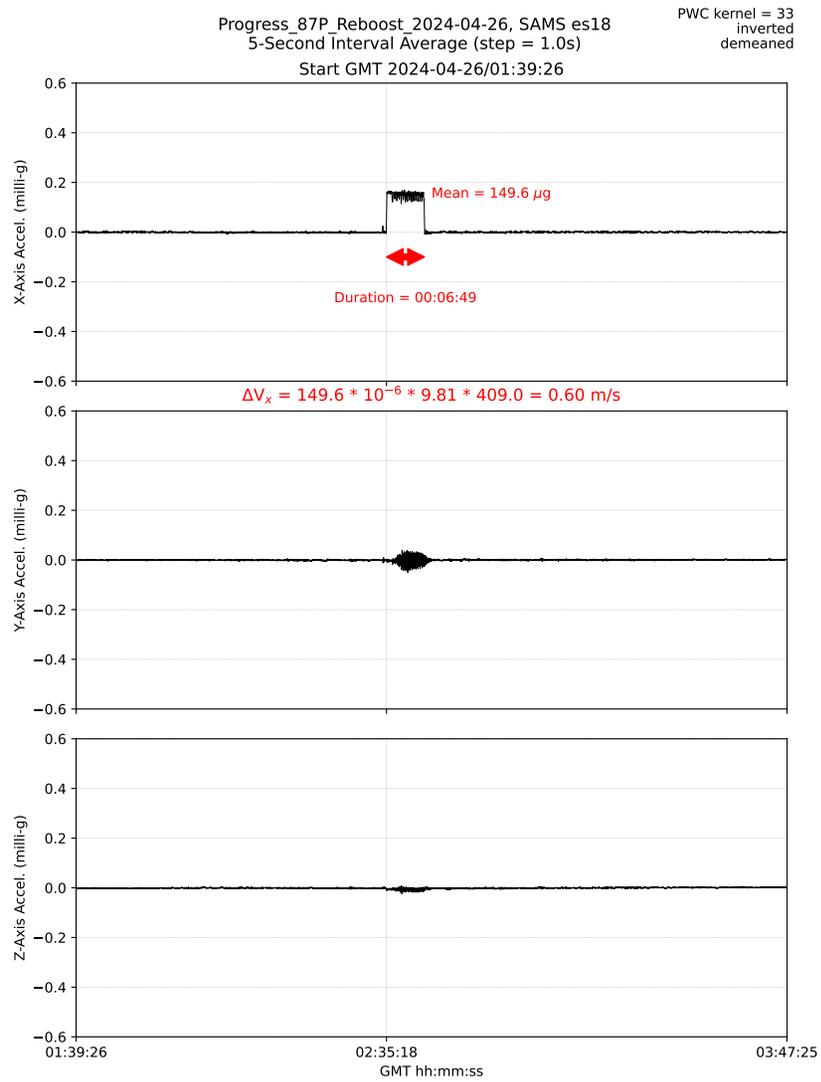


Fig. 4: 5-sec interval average for SAMS es18 sensor in the LAB.

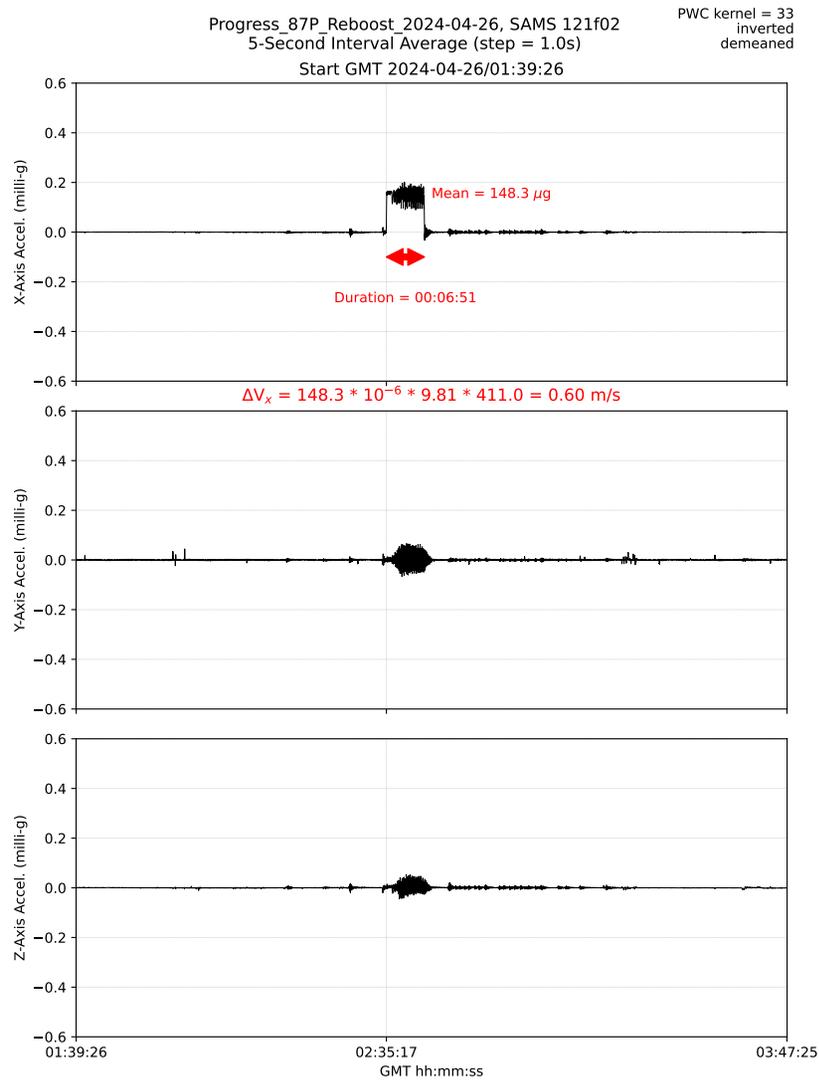


Fig. 5: 5-sec interval average for SAMS 121f02 sensor in the COL.

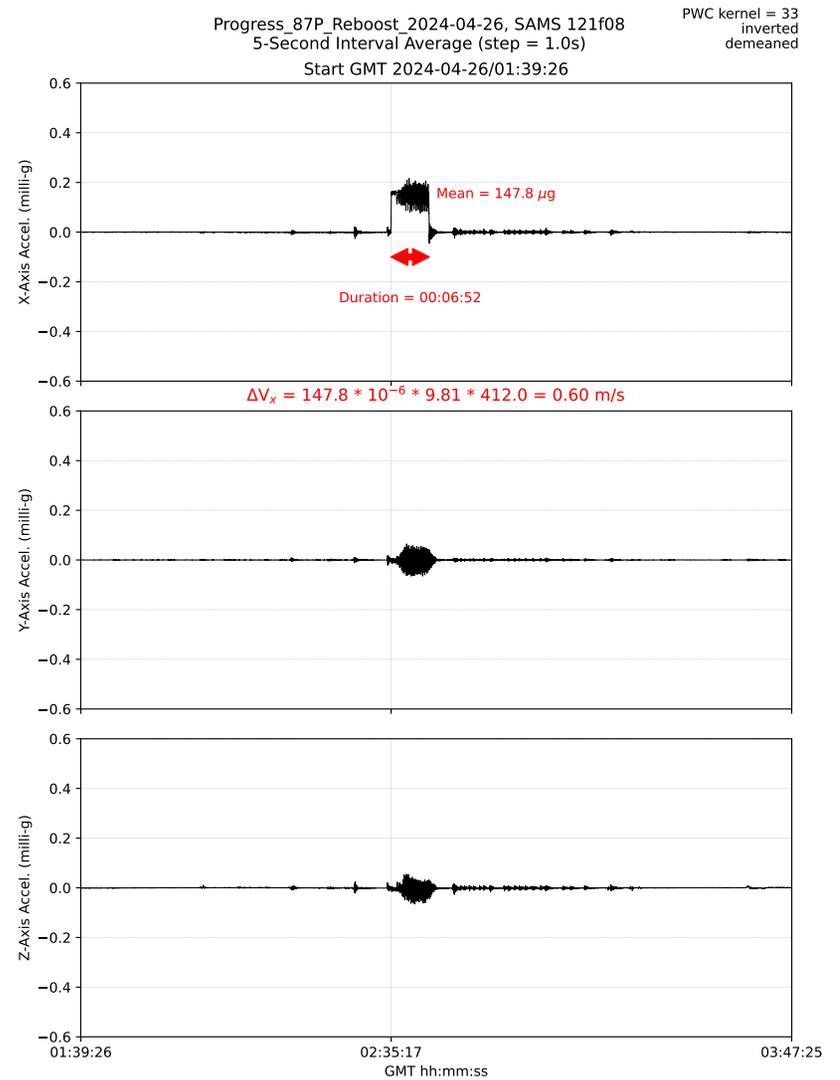


Fig. 6: 5-sec interval average for SAMS 121f08 sensor in the COL.

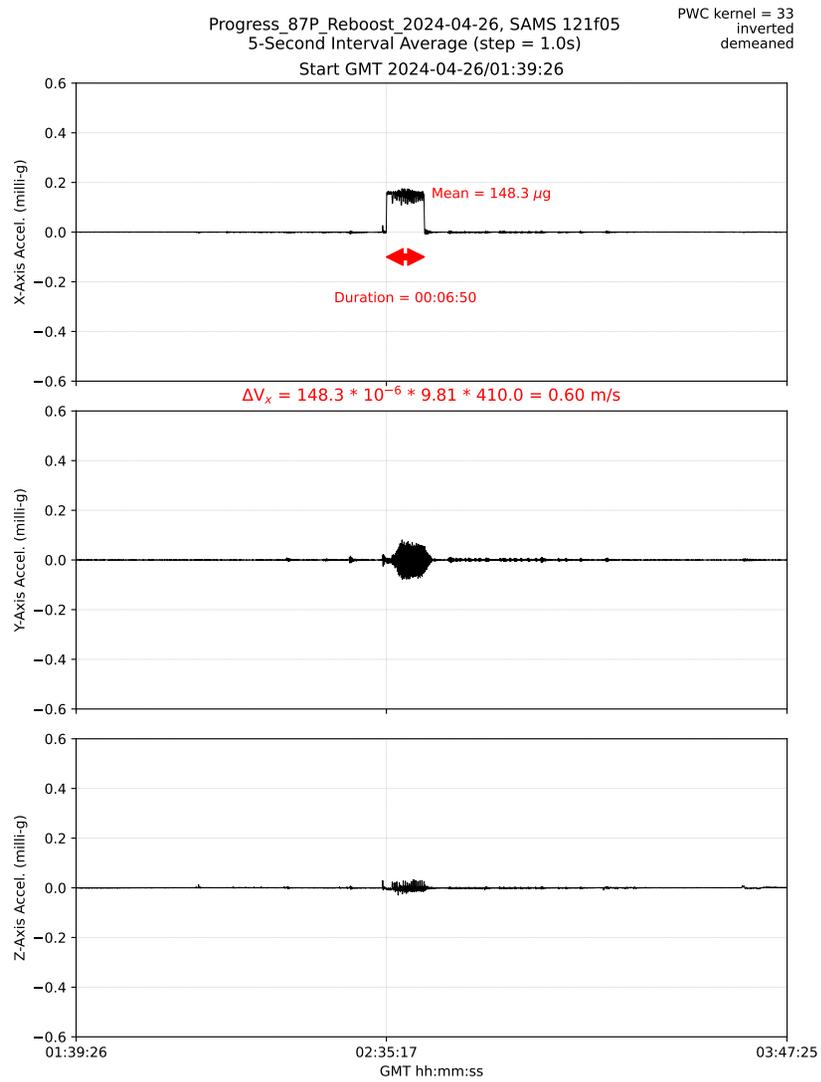


Fig. 7: 5-sec interval average for SAMS 121f05 sensor in the JEM.